

Supersymmetric Lepton Flavour Violation and Neutrinos mass textures: an update

This content has been downloaded from IOPscience. Please scroll down to see the full text.

2013 J. Phys.: Conf. Ser. 447 012059

(<http://iopscience.iop.org/1742-6596/447/1/012059>)

View [the table of contents for this issue](#), or go to the [journal homepage](#) for more

Download details:

IP Address: 150.214.168.144

This content was downloaded on 10/03/2015 at 08:43

Please note that [terms and conditions apply](#).

Supersymmetric Lepton Flavour Violation and Neutrinos mass textures: an update.

Mario E. Gómez, Mirco Cannoni

Department of Applied Physics, University of Huelva, 21071 Huelva, Spain

E-mail: mario.gomez@dfa.uhu.es, mirco.cannini@dfa.uhu.es

Abstract.

We revise the prospects for the observation of charged lepton flavour violation (LFV), in the light of the recent results from the LHC, MEG and neutrino experiments. We work in the context of the Minimal Supersymmetric Standard Model (MSSM) extended with massive neutrinos arising from a *see-saw* mechanism. The connection between the observed neutrino oscillations with flavour oscillations in the charged sector is established using a model for the Yukawa couplings arising from a SU(5) grand unified theory (GUT) with Abelian flavour symmetries. We discuss in this scenario the possible observation on the radiative decays $\mu \rightarrow e\gamma$ and $\tau \rightarrow \mu\gamma$ and LFV processes that could be detectable at a linear collider (LC) with a centre-of-mass energy in the TeV range.

1. Introduction

The picture of neutrino masses and oscillations has recently improved by the evidence for a non-zero value of θ_{13} as found by several reactor experiments [1, 2, 3]. The explanation for neutrino observations has to arise from theories beyond the standard model (SM). These theories also predict charged LFV. The link between neutrino oscillations and the violations of the individual lepton numbers $L_{e,\mu,\tau}$ offers the possibility of observing processes such as $l_i \rightarrow l_j\gamma$ ($i \neq j$) [4]. The present experimental upper limits, summarised below, already constrain significantly the parameter space of theoretical models,

$$BR(\mu \rightarrow e\gamma) < 5.7 \times 10^{-12} \quad [5], \quad (1)$$

$$BR(\tau \rightarrow \mu\gamma) < 4.4 \times 10^{-8} \quad [6], \quad (2)$$

$$BR(\tau \rightarrow e\gamma) < 3.3 \times 10^{-8} \quad [6]. \quad (3)$$

The strongest constraint on radiative decays is the latest MEG upper limit on $BR(\mu \rightarrow e\gamma)$ [5].

Supersymmetric (SUSY) theories connect the problem of SM fermions' flavour with the physics of their scalar partners. Experimental evidence on the absence or smallness of charged LFV phenomena seems to indicate that SUSY partners acquire masses with a flavour independent SUSY breaking mechanism. Models like the constrained minimal supersymmetric standard model (CMSSM) with universal scalar, gaugino masses and trilinear terms at the GUT scale extend the SM with an economical amount of undetermined parameters and without introducing flavour conflicts. We use CMSSM framework supplemented with the two elements that we need to explain neutrino oscillations: a structure of the Yukawa interactions and a



mechanism to explain small neutrino masses. As a model for the Yukawa textures we consider one inspired by SU(5) GUT models with Abelian flavour symmetries [7, 8] while tiny neutrino masses can be explained through a see-saw mechanism.

In the context of SUSY, even if the soft scalar masses were universal at the unification scale, quantum corrections between the GUT scale and low energies would modify this structure via renormalization-group running that generates off-diagonal contributions giving rise to observable LFV signals. Examples are rare lepton decays [4, 9, 10, 11, 12] and slepton production at a LC [13, 14, 15, 16].

We focus our analysis to models with universality assumptions of the CMSSM and that are compatible with searches at the LHC, and that also respect cosmological relic density considerations [17]. As an output of our analysis, we discuss detection prospects at a Linear Collider (LC), taking into account the constraints that the recent LHC supersymmetry searches and the measurement of the Higgs mass [18, 19] imposes on the observability of LFV processes.

In the following sections we make a brief survey of the Yukawa model we used and the constraints from LFV radiative decays before presenting our results for LC signals. Further details and a complete list of references are given in Ref [20].

2. Neutrino Mass Textures and Predictions for Neutrino Observables

We choose a model inspired by a SU(5) GUT combined with family symmetries [7, 8] such that the Yukawa textures take the form:

$$Y_\ell \propto \begin{pmatrix} \varepsilon^4 & \varepsilon^3 & \varepsilon \\ \varepsilon^3 & \varepsilon^2 & 1 \\ \varepsilon^3 & \varepsilon^2 & 1 \end{pmatrix}, \quad Y_\nu \propto \begin{pmatrix} \varepsilon^{|1\pm n_1|} & \varepsilon^{|1\pm n_2|} & \varepsilon^{|1\pm n_3|} \\ \varepsilon^{|n_1|} & \varepsilon^{|n_2|} & \varepsilon^{|n_3|} \\ \varepsilon^{|n_1|} & \varepsilon^{|n_2|} & \varepsilon^{|n_3|} \end{pmatrix} \quad (4)$$

where $Y_{\ell,\nu}$ stand for the Yukawa couplings of charged leptons and neutrinos respectively. n_i denote the U(1) charges and the heavy Majorana mass matrix is given by

$$M_N \propto \begin{pmatrix} \varepsilon^{2|n_1|} & \varepsilon^{|n_1+n_2|} & \varepsilon^{|n_1+n_3|} \\ \varepsilon^{|n_1+n_2|} & \varepsilon^{2|n_2|} & \varepsilon^{|n_2+n_3|} \\ \varepsilon^{|n_1+n_3|} & \varepsilon^{|n_2+n_3|} & \varepsilon^{2|n_3|} \end{pmatrix}. \quad (5)$$

The SU(5) structure of the model implies that the charged-lepton mass matrix is the transpose of the down-quark mass matrix, which relates the mixing of the left-handed leptons to that of the right-handed down-type quarks. The latest property implies that a large mixing can take place in the lepton sector while having a small mixing in the down quark sector, as suggested by a natural explanation of the Cabibbo-Kobayashi-Maskawa (CKM) matrix.

The light neutrinos mass matrix is, thanks to the see-saw mechanism,

$$m_{eff} \approx m_\nu^D \frac{1}{M_N} m_\nu^{D^T} \propto \begin{pmatrix} \varepsilon^2 & \varepsilon & \varepsilon \\ \varepsilon & 1 & 1 \\ \varepsilon & 1 & 1 \end{pmatrix}. \quad (6)$$

Note that $Y_\ell Y_\ell^\dagger \sim m_{eff}$ at the lowest order, thus, given the following diagonalizations of the Dirac and Majorana mass matrices,

$$V_\ell^T (Y_\ell Y_\ell^\dagger) V_\ell^* = \text{diag}(y_e^2, y_\mu^2, y_\tau^2), \quad (7)$$

$$V_D^T (Y_\nu Y_\nu^\dagger) V_D^* = \text{diag}(y_{\nu_1}^2, y_{\nu_2}^2, y_{\nu_3}^2), \quad (8)$$

$$U_N^T M_N U_N = \text{diag}(M_1, M_2, M_3), \quad (9)$$

$$U_\nu^T m_{eff} U_\nu = \text{diag}(m_{\nu_1}, m_{\nu_2}, m_{\nu_3}), \quad (10)$$

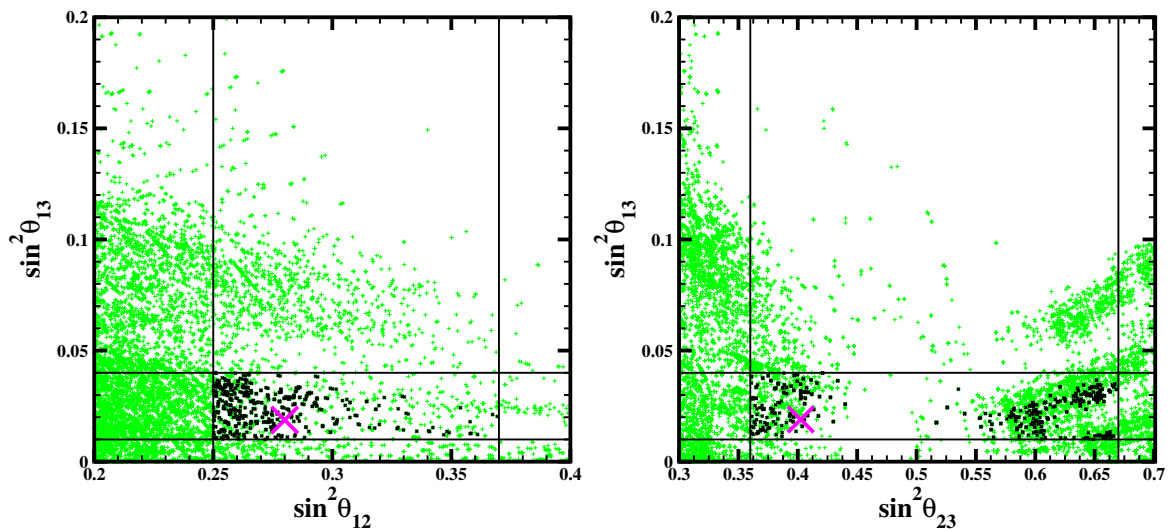


Figure 1. Correlations between the neutrino mixing angles before and after imposing the constraints on the model coefficients as discussed in the text. The solid lines indicate the experimental bounds, and the small black crosses represent models satisfying all the constraints. The large magenta cross corresponds to model that is discussed in the text and used for numerical calculations.

V_ℓ and U_ν diagonalize matrices with a similar structure. The Maki-Nakagawa-Sakata (MNS) matrix is given by

$$U_{MNS} \equiv U = V_\ell^\dagger U_\nu, \quad (11)$$

The texture is determined up to coefficients $\mathcal{O}(1)$ in every entry of the matrices. We take the expansion parameter to be $\varepsilon = 0.2$ and multiply the entries of Y_ℓ , Y_ν and M_N given by Eqs. (4) and (5) by coefficients in the range $\pm[0.5, 2]$. We require the ranges for the mixing angles of Ref. [21] and assume a neutrino mass hierarchy with a mass splitting of the order of the neutrino masses and $m_{\nu_3} \sim 0.05$ eV. In addition we make a further selection by requiring:

- The coefficients of Y_ν and M_N are matched to a light neutrino mass matrix m_{eff} of the form (6), with entries that deviate by a factor between 1/2 and 2 from those in Eq. (6).
- That the hierarchy of eigenvalues of $Y_\ell Y_\ell^\dagger$ (which, as discussed above, has a similar structure to m_{eff} and $Y_\nu Y_\nu^\dagger$) preserves the order of the gauge eigenstates.

After a search of about 10^9 fits we display in Fig. 1 the predictions for the neutrino mixing angles corresponding to the above criteria. We can see that, within this class of models, most of solutions that reproduce the correct range of θ_{12} and θ_{23} also predict a neutrino mixing angle θ_{13} that is compatible with the data from [1, 2] and also with the global analysis of neutrino data [22]. We find values for θ_{12} mostly on the lower part of its experimental range, while values with maximal $\theta_{23} = \pi/4$ are not frequent. This would have been the case when U_{MNS} arose from a m_{eff} with the structure of Eq. (6) and moderate 2-3 mixing from the charged lepton sector. However, with our conditions, both V_ℓ and U_ν in Eq. (11) arise from diagonalization of matrices with a large 2-3 mixing and our solutions do not present a distribution of $\sin \theta_{23}$ around 1/2.

We selected the marked point as reference for further analysis, this corresponds to:

$$Y_\ell \propto \begin{pmatrix} \varepsilon^4 & 2\varepsilon^3 & -1.75\varepsilon \\ -0.5\varepsilon^3 & 1.9\varepsilon^2 & 0.5 \\ -0.5\varepsilon^3 & -0.7\varepsilon^2 & 1.25 \end{pmatrix}, \quad Y_\nu \propto \begin{pmatrix} \varepsilon^{|1\pm n_1|} & \varepsilon^{|1\pm n_2|} & 2\varepsilon^{|1\pm n_3|} \\ 0.75\varepsilon^{|n_1|} & \varepsilon^{|n_2|} & -0.5\varepsilon^{|n_3|} \\ \varepsilon^{|n_1|} & \varepsilon^{|n_2|} & 1.25\varepsilon^{|n_3|} \end{pmatrix},$$

$$M_N \propto \begin{pmatrix} \varepsilon^{2|n_1|} & \varepsilon^{|n_1+n_2|} & -\varepsilon^{|n_1+n_3|} \\ \varepsilon^{|n_1+n_2|} & \varepsilon^{2|n_2|} & \varepsilon^{|n_2+n_3|} \\ -\varepsilon^{|n_1+n_3|} & \varepsilon^{|n_2+n_3|} & -\varepsilon^{2|n_3|} \end{pmatrix}. \quad (12)$$

The predictions for the neutrino angles are, $\sin^2 \theta_{13} = 0.019$, $\sin^2 \theta_{12} = 0.28$, $\sin^2 \theta_{23} = 0.40$. We observe no correlation between LFV and particular arrangements of neutrino parameters. However, LFV is maximized by large out-diagonal elements on Y_ℓ and by the choice of charges n_i . Neutrino fits are independent of these charges because despite they affect Y_ν and M_N but not their combination in m_{eff} .

3. Scenarios for Charged-Lepton-Flavour Violation

A complete GUT model with an Abelian symmetry that determines the Yukawa textures also implies a flavour dependence of the soft mass matrices on the family charges. The determination of the resulting model at M_{GUT} depends on the details of the SUSY breaking and RGE evolution of the soft terms [23]. In this work we assume that this model is the MSSM extended by a see-saw with soft terms that do not deviate significantly from the universality conditions of the CMSSM.

Even if we start with universal soft-terms at M_{GUT} , at the intermediate scale where the see-saw takes place, M_3 (taken as the mass of the heaviest Majorana neutrino), the slepton mass matrices and Y_ℓ cannot be diagonalized with a single superfield rotation. Thus, the interactions lepton-slepton-gaugino can mix flavours. To understand this mismatch of the leptons and sleptons rotations we can consider the soft masses evolution from M_{GUT} to M_3 in a basis such that Y_ν is diagonal, at M_3 the right handed neutrinos decouple and the REG can be rewritten in terms of Y_ℓ diagonal. In this basis the soft terms involving left slepton are not diagonal and can be written in terms of the matrix $V_{LFV} = V_D^\dagger V_\ell$:

$$m_{LL}^2 = V_{LFV}^\dagger (m_{LL}^2)_d V_{LFV}, \quad (13)$$

while the A -terms become:

$$A_\ell = V_{LFV}^T (A_\ell)_d. \quad (14)$$

Here $(m_{LL}^2)_d$ and $(A_\ell)_d$ are the soft terms resulting from the RGE running of the universal soft terms at the GUT scale to M_3 with the fields written in a basis such that Y_ν is diagonal. The successive run of the MSSM RGEs to M_{SUSY} do not have a significant effect on the off-diagonal elements.

The matrix V_{LFV} can be evaluated from Yukawa textures. We use the fit of Eq. (12) obtained by adjusting the matrices to the neutrino parameters at a low energy scale. RGE effects are taken into account by considering the global scale evolution of the neutrino masses set initially by imposing $m_{\nu_3} = 0.05$ eV and the see-saw condition at a common scale M_3 to determine the strength of the largest Dirac Yukawa coupling. Since we assume hierarchical neutrinos, we neglect the evolution of the non-diagonal matrix elements of effective neutrino mass and the lightest generations. These effects are small [24] and can be included in the uncertainty of the coefficients without affecting the stability of the solutions. Furthermore, no sizeable effect is expected from including lower right handed neutrino scales since the Dirac Yukawa decreases

Table 1. Values of V_{LFV} with $\varepsilon = 0.2$ for the fit of Eq. 12.

n_i	$c_1 = \{n_1 = 2, n_2 = 0, n_3 = 1\}$	$c_2 = \{n_1 = 0, n_2 = 1, n_3 = 0\}$
V_{LFV}	$\begin{pmatrix} -0.820 & 0.343 & 0.458 \\ 0.494 & 0.829 & 0.263 \\ -0.289 & 0.441 & -0.849 \end{pmatrix}$	$\begin{pmatrix} 0.806 & -0.401 & -0.436 \\ -0.437 & -0.899 & 0.016 \\ -0.399 & 0.178 & -0.901 \end{pmatrix}$

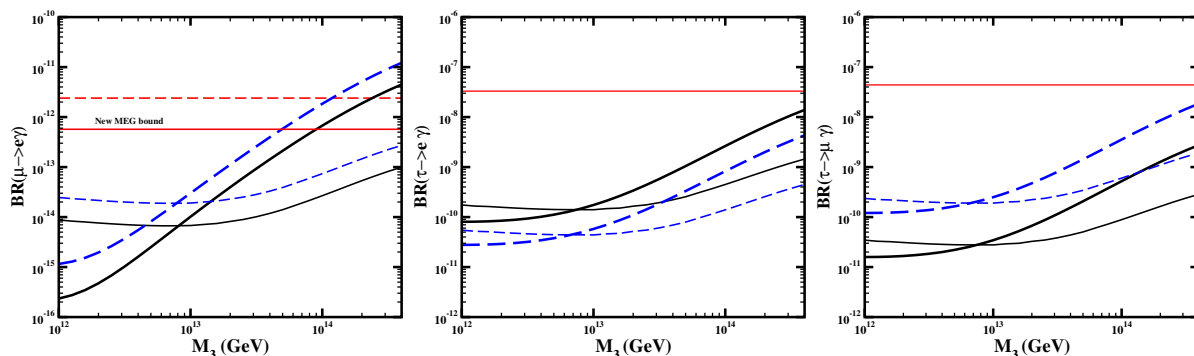


Figure 2. Predictions for the rare LFV decays $\ell_i \rightarrow \ell_j \gamma$ as a function of the heaviest right-handed neutrino mass M_N , for the benchmark point displayed in (15) (a) (thick line), (b) (thin line) and using some the neutrino mixing fits shown in Table 1. The solid lines correspond to cases (c_1), dashed to (c_2). The horizontal solid lines indicate the current experimental upper bounds.

with them becoming tiny when M_1 and M_2 are much lower than M_3 . In our computation we use the matrices given in Eq. (12), V_ℓ has large out-diagonal elements and the n_i charges that are not relevant to fit neutrino data become important in the determination of LFV effects as can be seen in the values of V_{LFV} of table 1. We need now to evaluate the SUSY spectrum and couplings using the CMSSM as general framework. The recent LHC measurement of the Higgs mass seems to point towards large SUSY masses in these scenarios. However, the global analysis of the CMSSM parameter space of Ref. [17] yielded two almost equally good fits to the available data, one with relatively light sparticle masses and $\tan \beta \sim 16$, and the other with heavier sparticle masses and $\tan \beta \sim 45$, corresponding to the following CMSSM sets of parameters:

$$\begin{aligned}
 (a) \quad & \tan \beta = 16, \quad m_0 = 300 \text{ GeV}, \quad M_{1/2} = 910 \text{ GeV}, \quad A_0 = 1320 \text{ GeV}, \\
 (b) \quad & \tan \beta = 45, \quad m_0 = 1070 \text{ GeV}, \quad M_{1/2} = 1890 \text{ GeV}, \quad A_0 = 1020 \text{ GeV}. \quad (15)
 \end{aligned}$$

Point (a) belongs to the region where the WMAP-favoured range of $\Omega_\chi h^2$ is achieved via $\chi - \tilde{\tau}$ coannihilation. Point (b) lies in the funnel region where the neutralino LSP annihilates rapidly via direct-channel H/A poles.

4. LFV in radiative decays.

The branching ratios (BRs) of the decays $\ell_j \rightarrow \ell_i + \gamma$ are calculated using full mass matrices diagonalization and the formulae of Ref. [9]. In Fig. 2 we show numerical predictions for the LFV branching ratios. We can see the effect of varying M_N from 6×10^{14} GeV down to 10^{12}

GeV for the choices of right-handed neutrino charges of Table 1. The branching ratios are larger for the lower-mass scenario with $\tan\beta = 16$, due to the lighter spectrum. The new MEG bound on $\text{BR}(\tau \rightarrow \mu\gamma)$ imposes constraints on the "see-saw" scale for all for all the charge choices of Table 1 at point (a) while the predictions of point (b) are in the range of the experimental searches.

5. LFV Observation at a Linear Collider

In Ref. [13] it was shown that if the flavour mixing is introduced in the LL slepton sector, as is the case for the models under consideration here, slepton-pair production and LFV decays such as:

$$\begin{aligned} e^+e^- &\rightarrow \tilde{\ell}_i^- \tilde{\ell}_j^+ \rightarrow \tau^\pm \mu^\mp \tilde{\chi}_1^0 \tilde{\chi}_1^0, \\ e^+e^- &\rightarrow \tilde{\nu}_i \tilde{\nu}_j^c \rightarrow \tau^\pm \mu^\mp \tilde{\chi}_1^+ \tilde{\chi}_1^- \end{aligned} \quad (16)$$

lead to a cross section of the order of 1 fb, the reference value used in [13], for a LFV signal of $\mu^\pm\tau^\pm$ pairs that can be distinguished from the background, according to the considerations in [15].

Here, we extend our previous results by considering the full structure of the Yukawa couplings, thus comparing the LFV production of charged leptons of all generations and use the updated supersymmetric benchmark points compatible with the LHC measurement of m_h .

We present in Fig. 3 the expected cross sections as functions of \sqrt{s} for the same sets of charges as those in Fig. 2. For definiteness, the right-handed neutrino mass scale is set to $M_N = 10^{14}$ GeV for point (b) and the restrictive value of $M_N = 2 \times 10^{13}$ such that the new MEG bound on $\text{BR}(\mu \rightarrow e\gamma)$ is preserved. We observe that, by varying this scale in a similar range as in Fig. 2, the cross sections also change by approximately an order of magnitude: according to the see-saw conditions, larger values of M_N imply larger Dirac neutrino Yukawa couplings, and thus stronger RGE effect. The results depend sensitively on the sparticle spectrum. If we restrict ourselves to the benchmark CMSSM parameter choice discussed above, we would require \sqrt{s} at TeV scale in order to have significant LFV signals in slepton mixing.

6. Conclusions

In our presentation, we revisited the signatures of charged LFV in theoretically-motivated scenarios, studying the correlations arising in CMSSM models with parameter values that are favoured by the LHC and cosmological considerations. We have explored these issues using updated experimental input from neutrino data, particularly recent measurements of θ_{13} , MEG and the LHC.

Models based in SU(5) with Abelian flavour symmetries provide very interesting possibilities for understanding the hierarchy of fermion masses and mixings. We performed a big scan of fits to the neutrino data paying attention to the the naturalness of the fit and avoiding artificial cancellations arising from specific choices of coefficients. We get a pattern of neutrino predictions and correlations compatible with the global analysis of neutrino data [22]. In general, we found that fittings with similar predictions for the neutrino parameters may lead to very different LFV predictions. This can provide information on the heavy Majorana neutrino matrix.

The new MEG bound on $\mu \rightarrow e\gamma$ is still compatible with the LHC data. At the point with lighter SUSY spectrum it determines an upper limit for the see-saw scale while for point with large SUSY spectrum, LFV radiative decays are still interesting for experimental searches. Regarding LC prospects, it was possible to establish correlations between the expected rates for radiative LFV decays and LFV $\tau\mu$ pair production at a future LC. Within the CMSSM framework studied, the absence of a supersymmetry signal in the LHC data and the discovery

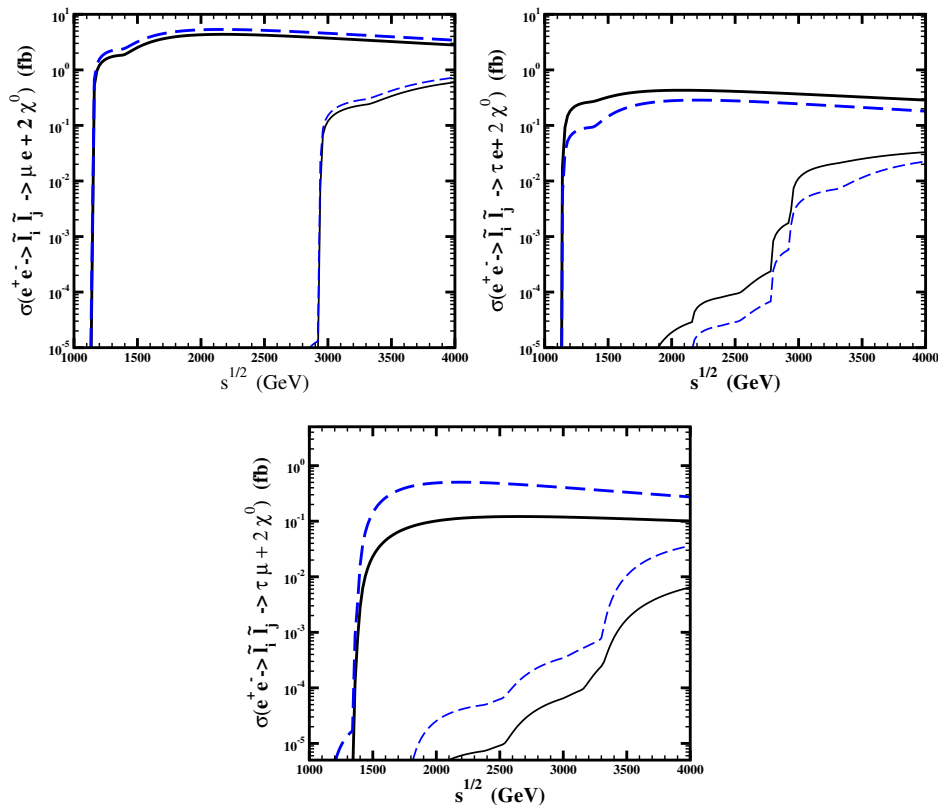


Figure 3. Values of the cross sections $\sigma(e^+e^- \rightarrow \tilde{l}_i^- \tilde{l}_j^+ \rightarrow \ell_a^\pm \ell_b^\mp + 2\chi^0)$ ($\ell_a \neq \ell_b$ as indicated for each panel) as functions of \sqrt{s} . The line styles are the same as those in Fig. 2. For the point (a) we use $M_N = 2 \times 10^{13}$ GeV. While for the point (b) we use $M_N = 10^{14}$ GeV.

of a neutral Higgs weighing ~ 125 GeV imply that observation of slepton flavour violation at a LC will be possible for only energies beyond 1 TeV.

Acknowledgments. The work of M. C. is supported by MultiDark under Grant No. CSD2009-00064 of the Spanish MICINN Consolider-Ingenio 2010 Program. M. E. G. and M. C. acknowledge further support from the MICINN project FPA2011-23781, from the Grant MICINN-INFN(PG21)AIC-D-2011-0724 and the project P07FQM02962 funded by “Junta de Andalucía”.

- [1] F. P. An *et al.* [Daya Bay Collaboration], Phys. Rev. Lett. **108** (2012) 171803.
- [2] J. K. Ahn *et al.* [RENO Collaboration], Phys. Rev. Lett. **108** (2012) 191802.
- [3] Y. Abe *et al.* [Double Chooz Collaboration], arXiv:1301.2948 [hep-ex].
- [4] For reviews, see:
M. Raidal *et al.*, Eur. Phys. J. C **57** (2008) 13; Y. Kuno and Y. Okada, Rev. Mod. Phys. **73** (2001) 151.
- [5] J. Adam *et al.* [MEG Collaboration], arXiv:1303.0754 [hep-ex].
- [6] K. Nakamura *et al.* [Particle Data Group], J. Phys. G **37** (2010) 075021.
- [7] S. Lola and G.G. Ross, Nucl. Phys. B **553** (1999) 81.
- [8] J. R. Ellis, M. E. Gómez and S. Lola, JHEP **0707** (2007) 052.
- [9] J. Hisano, T. Moroi, K. Tobe and M. Yamaguchi, Phys. Rev. D **53** (1996) 2442.
- [10] M. Gómez, G. Leontaris, S. Lola and J. Vergados, Phys. Rev. **D 59** (1999) 116009.
- [11] J. R. Ellis, M. E. Gómez, G. K. Leontaris, S. Lola and D. V. Nanopoulos, Eur. Phys. J. C **14** (2000) 319.
- [12] S. Antusch, E. Arganda, M. J. Herrero and A. M. Teixeira, JHEP **0611** (2006) 090 [hep-ph/0607263].
- [13] E. Carquin, J. Ellis, M. E. Gómez and S. Lola, JHEP **1111** (2011) 050.

- [14] J. Hisano, M. M. Nojiri, Y. Shimizu and M. Tanaka, Phys. Rev. D **60**, 055008 (1999).
- [15] F. Deppisch, H. Päs, A. Redelbach, R. Rückl and Y. Shimizu, Phys. Rev. D **69** (2004) 054014.
- [16] A. Abada, A. J. R. Figueiredo, J. C. Romão and A. M. Teixeira, JHEP **1208** (2012) 138 [arXiv:1206.2306 [hep-ph]].
- [17] O. Buchmueller *et al.*, Eur. Phys. J. C **72** (2012) 2243 [arXiv:1207.7315 [hep-ph]].
- [18] G. Aad *et al.* [ATLAS Collaboration], Phys. Lett. B **716** (2012) 1.
- [19] S. Chatrchyan *et al.* [CMS Collaboration], Phys. Lett. B **716** (2012) 30.
- [20] M. Cannoni, J. Ellis, M. E. Gomez, S. Lola and , arXiv:1301.6002 [hep-ph].
- [21] M. C. González-García, M. Maltoni, J. Salvado and T. Schwetz, arXiv:1209.3023 [hep-ph].
- [22] G.L. Fogli, E. Lisi, A. Marrone, D. Montanino and A. Palazzo, A.M. Rotunno, Phys. Rev. D **86** (2012) 013012.
- [23] M.E. Gómez, S. Lola, P. Naranjo and J. Rodríguez-Quintero, JHEP **0904** (2009) 043.
- [24] J. R. Ellis, S. Lola and , Phys. Lett. B **458** (1999) 310.